

# Lecture 3: Image Restoration

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B14 Image Analysis

Michaelmas 2014

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- Image degradations
  - motion blur, focus blur, resolution
- The inverse filter
- The Wiener filter
- MAP formulation

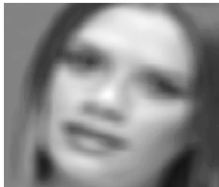
In contrast to image enhancement, in image restoration the degradation is **modelled**. This enables the effects of the degradation to be (largely) removed

# Degradations

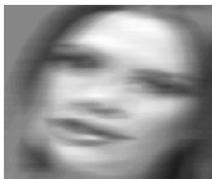
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- original



- optical blur



- motion blur



- spatial quantization (discrete pixels)



- additive intensity noise

# Overview – Deconvolution

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The objective is to restore a degraded image to its original form.

An observed image can often be modelled as:

$$g(x, y) = \int \int h(x - x', y - y') f(x', y') dx' dy' + n(x, y)$$

where the integral is a convolution,  $h$  is the point spread function of the imaging system, and  $n$  is additive noise.

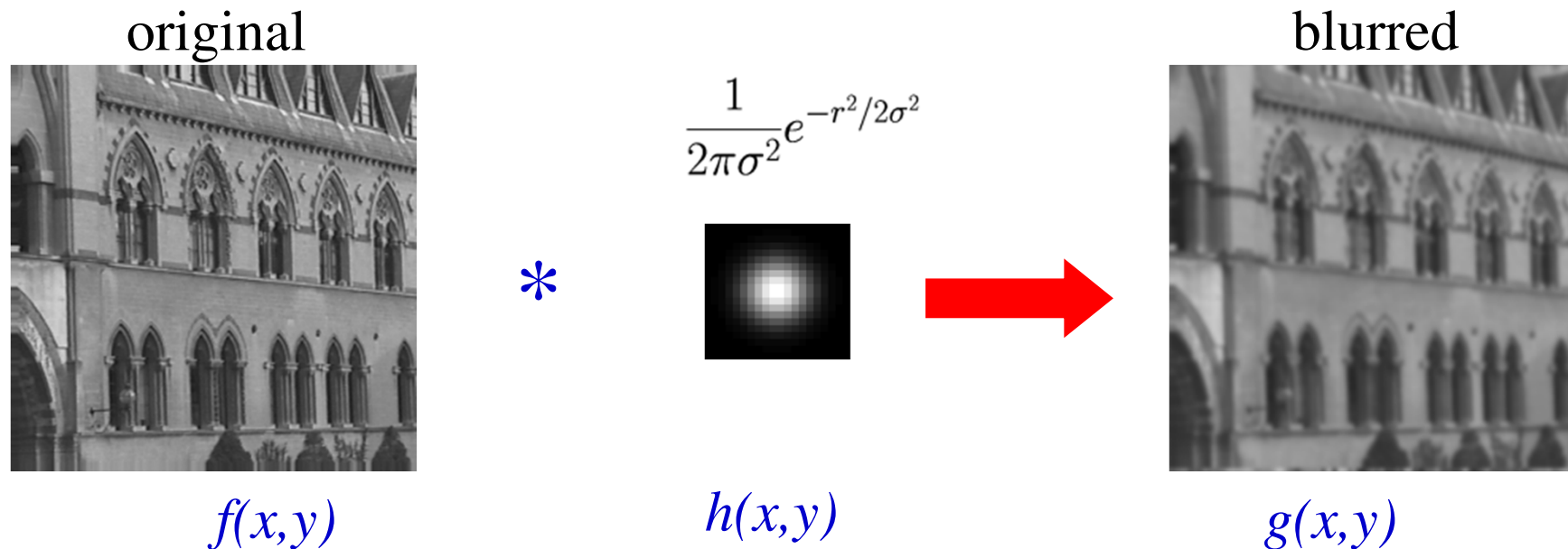
The objective of image restoration in this case is to estimate the original image  $f$  from the observed degraded image  $g$ .

# Degradation model

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Model degradation as a convolution with a linear, shift invariant, filter  $h(x,y)$

- Example: for out of focus blurring, model  $h(x,y)$  as a Gaussian

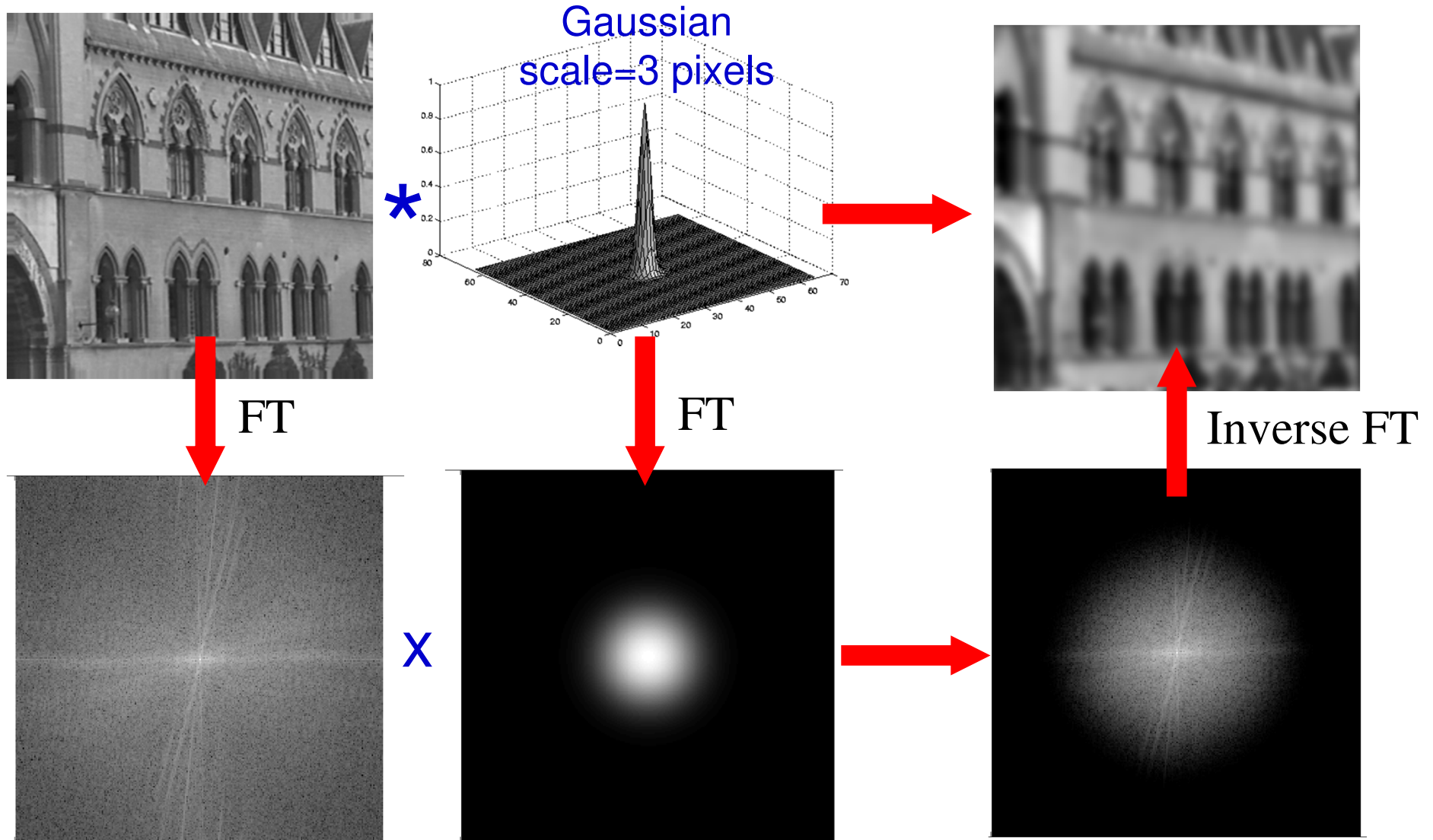


i.e. :  $g(x,y) = h(x,y) * f(x,y)$

$h(x,y)$  is the impulse response or point spread function of the imaging system



# The challenge: loss of information and noise

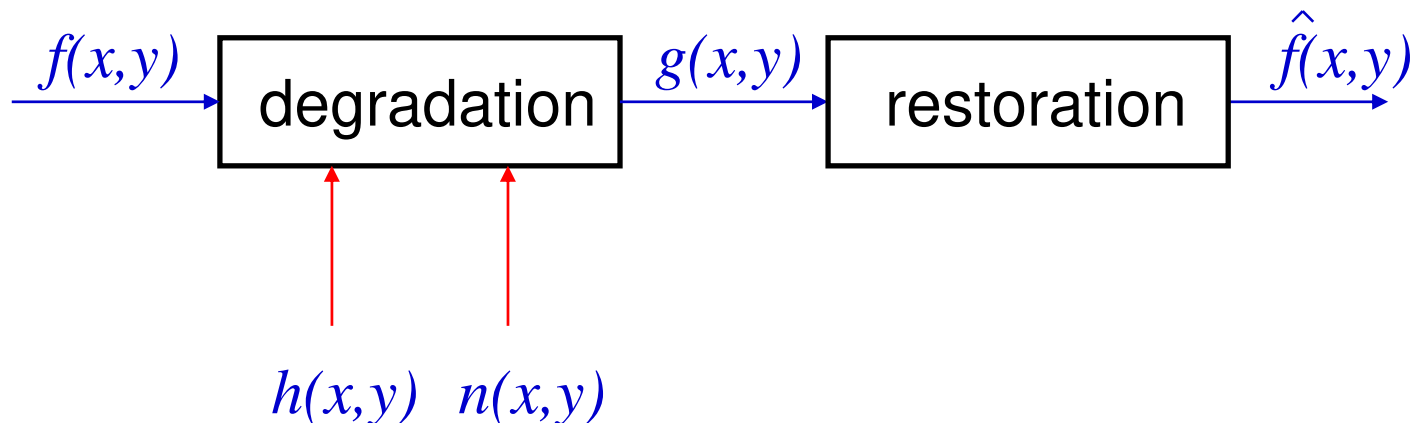


Blurring acts as a low pass filter and attenuates higher spatial frequencies

# Definitions

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- $f(x,y)$  – image before degradation, ‘true image’
- $g(x,y)$  – image after degradation, ‘observed image’
- $h(x,y)$  – degradation filter
- $\hat{f}(x,y)$  – estimate of  $f(x,y)$  computed from  $g(x,y)$
- $n(x,y)$  – additive noise



$$g(x,y) = h(x,y) * f(x,y) + n(x,y) \Leftrightarrow G(u,v) = H(u,v) F(u,v) + N(u,v)$$

# The inverse filter

Start from the **generative** model

$$g(x,y) = h(x,y)*f(x,y) + n(x,y) \Leftrightarrow G(u,v) = H(u,v) F(u,v) + N(u,v)$$

and for the moment ignore  $n(x,y)$ , then an estimate of  $f(x,y)$  is obtained from

$$\hat{F}(u,v) = G(u,v) / H(u,v)$$

Restoration with an **inverse** filter



# 1D vector explanation

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$$g = h * f$$

$$g_n = \frac{1}{4}f_{n-1} + \frac{1}{2}f_n + \frac{1}{4}f_{n+1}$$

$$g_1 = \frac{1}{4}f_0 + \frac{1}{2}f_1 + \frac{1}{4}f_2$$

$$g_2 = \frac{1}{4}f_1 + \frac{1}{2}f_2 + \frac{1}{4}f_3$$

For  $n$  points (ignoring boundaries)

$$\mathbf{g} = \mathbf{A}\mathbf{f}$$

where  $\mathbf{g}$  and  $\mathbf{f}$  are  $n$ -vectors, and  $\mathbf{A}$  is an  $n \times n$  matrix. Hence, ignoring possible problems if  $\mathbf{A}$  is singular,

$$\mathbf{f} = \mathbf{A}^{-1}\mathbf{g}$$

## Fourier trick

$$g = h * f$$

$$G = HF$$

$$F = G/H$$

$$f = \text{IFT}(G/H)$$

## Example : Deblurring (deconvolution)

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Image blurred with Gaussian point spread function

$$h(x,y) = \frac{1}{2\pi\sigma^2} e^{-r^2/2\sigma^2}$$

$n(x,y)$  = Normal distribution, mean zero



$f(x,y)$



$g(x,y)$

blur  $\sigma = 1.0$  pixels

noise  $\sigma = 0.3$  grey levels

Restoration with an **inverse** filter

$\hat{F}(u,v) = G(u,v) / H(u,v)$  where  $H(u,v)$ , is the FT of the Gaussian

# Deblurring with an inverse filter

noise  $\sigma = 0.3$  grey levels

$$\hat{F}(u,v) = G(u,v) / H(u,v)$$

blur  $\sigma = 0.5$  pixels

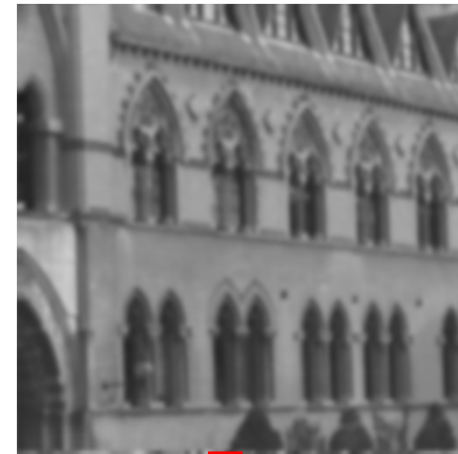


$g(x,y)$

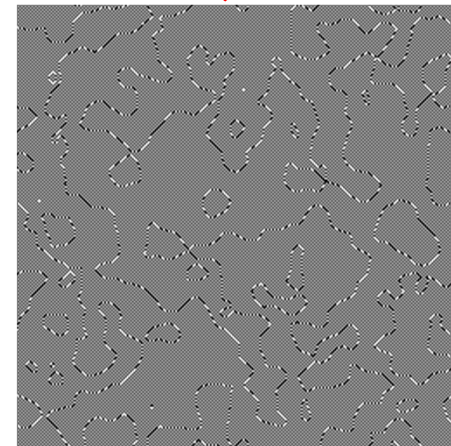
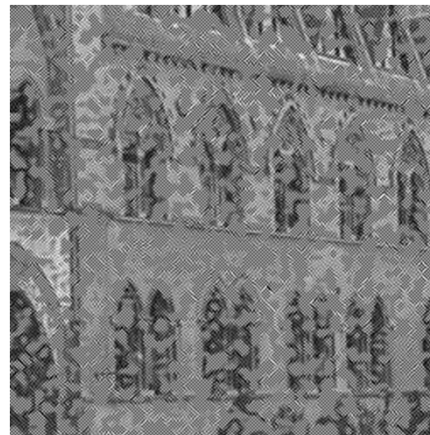
blur  $\sigma = 1.0$  pixels



blur  $\sigma = 1.5$  pixels



$\hat{f}(x,y)$



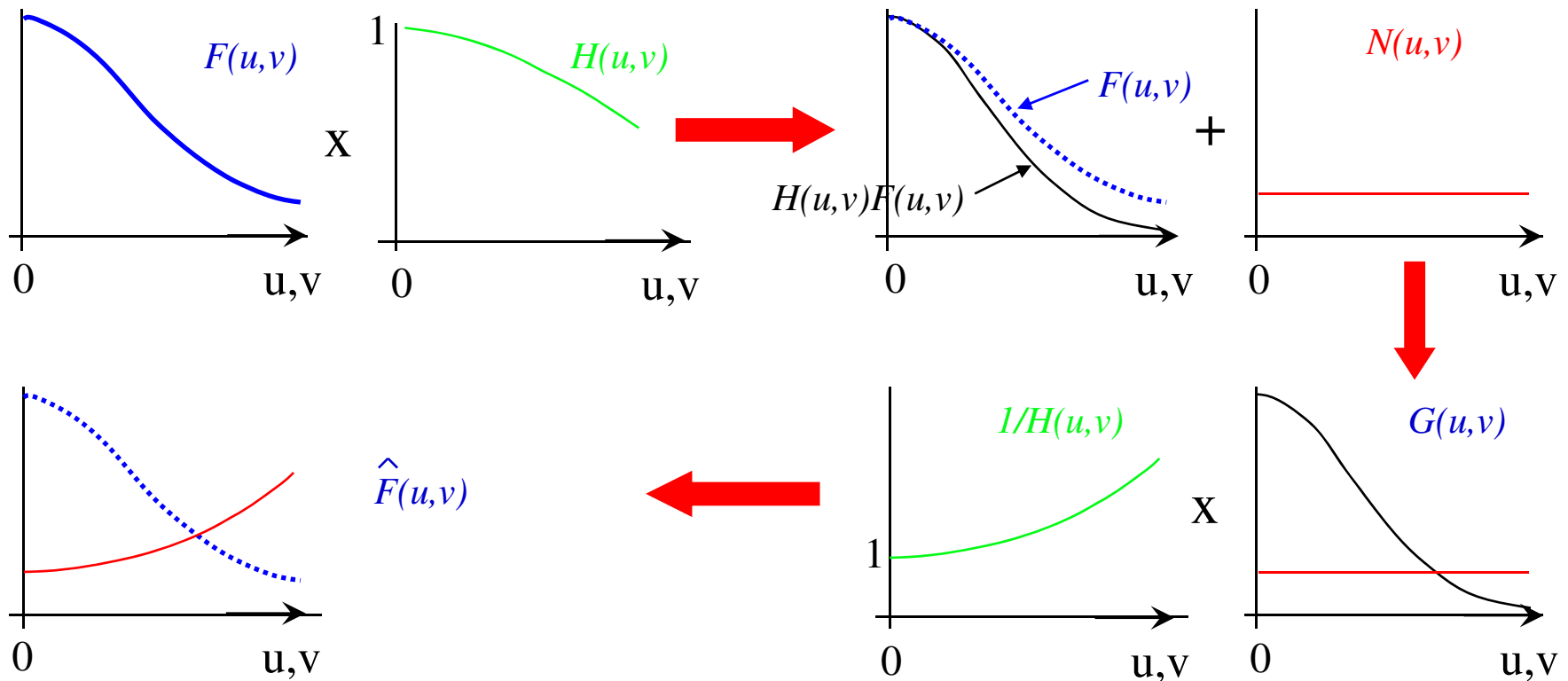
# The problem of noise amplification

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$$G(u,v) = H(u,v) F(u,v) + N(u,v)$$

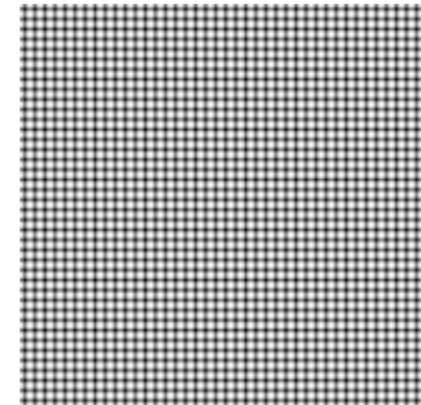
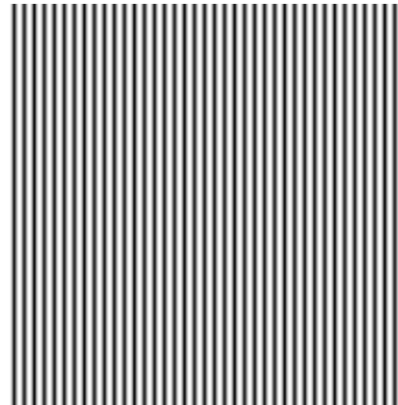
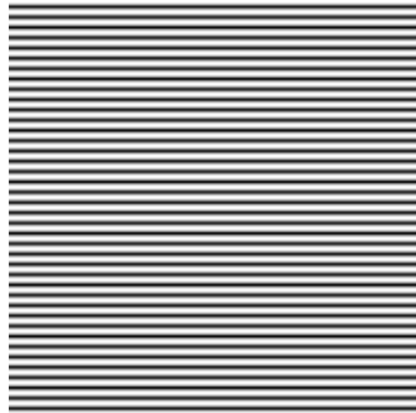
$$\hat{F}(u,v) = G(u,v) / H(u,v) = F(u,v) + N(u,v) / H(u,v)$$

Schematically ...



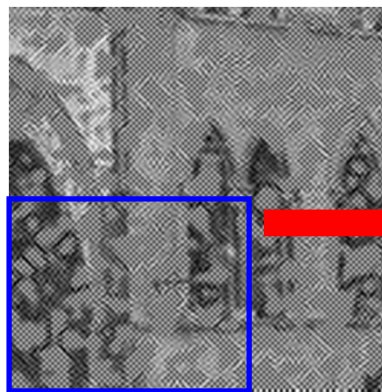
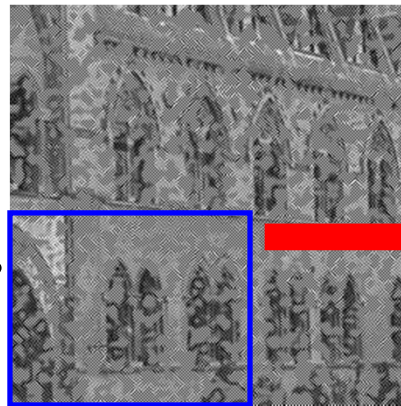


high spatial frequency sinusoids



$$\hat{f}(x,y)$$

blur  $\sigma = 1.0$  pixels



# The Wiener filter

# The Wiener filter

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$$\hat{F}(u, v) = W(u, v) G(u, v)$$

$$W(u, v) = \frac{H^*(u, v)}{|H(u, v)|^2 + K(u, v)}$$

where

$$K(u, v) = S_\eta(u, v) / S_f(u, v)$$

$$S_f(u, v) = |F(u, v)|^2 \text{ power spectral density of } f(x, y)$$

$$S_\eta(u, v) = |N(u, v)|^2 \text{ power spectral density of } \eta(x, y)$$

# Frequency behaviour

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$$\hat{F}(u, v) = W(u, v) G(u, v)$$

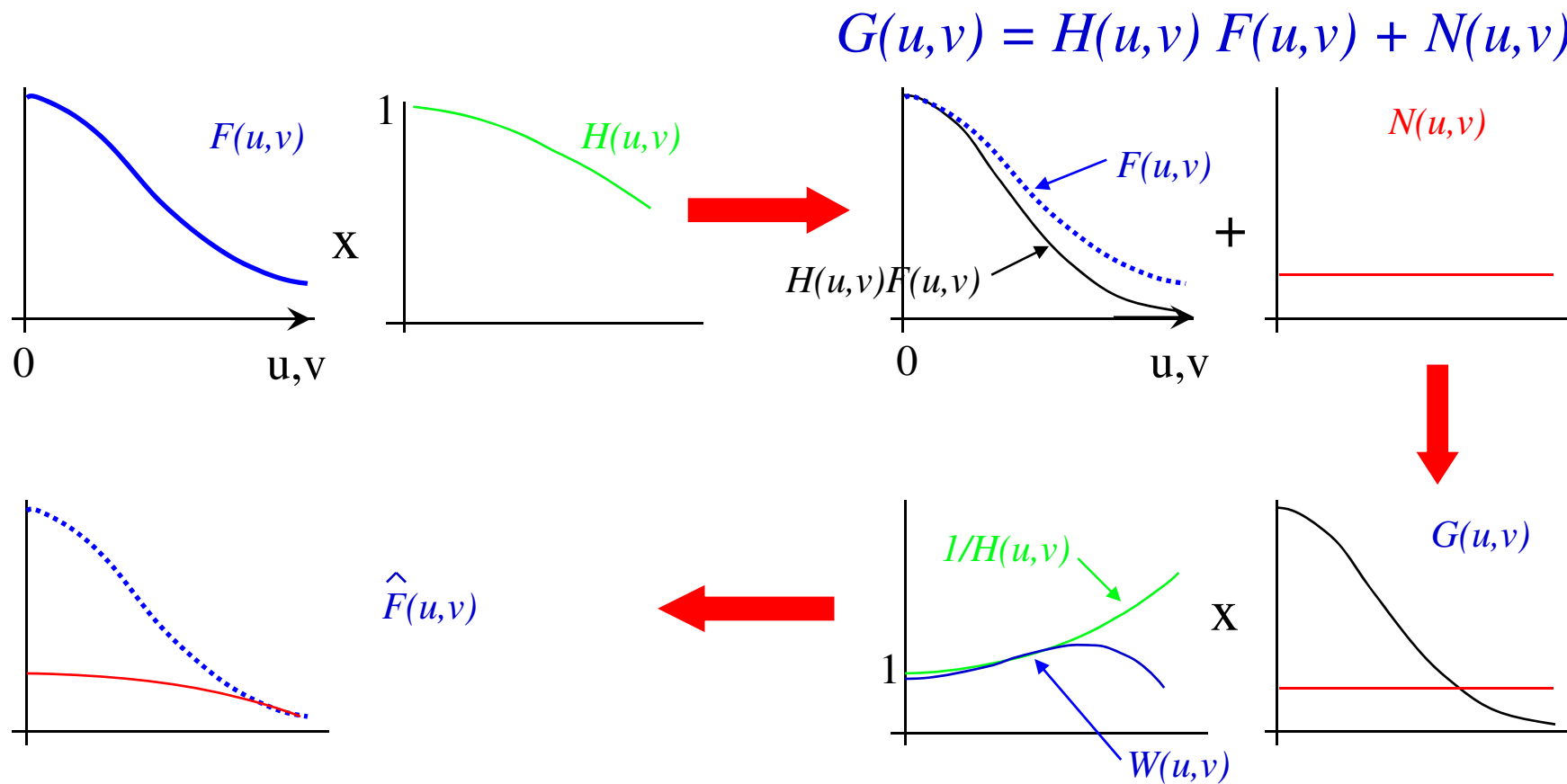
$$W(u, v) = \frac{H^*(u, v)}{|H(u, v)|^2 + K(u, v)}$$

- If  $K = 0$  then  $W(u, v) = 1 / H(u, v)$ , i.e. an inverse filter
- If  $K \gg |H(u, v)|$  for large  $u, v$ , then high frequencies are attenuated
- $|F(u, v)|$  and  $|N(u, v)|$  are often known approximately, or
- $K$  is set to a constant scalar which is determined empirically
- A Wiener filter minimizes the least square error  $\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} (f(x, y) - \hat{f}(x, y))^2 dx dy$

$$\hat{F}(u,v) = W(u,v) G(u,v)$$

$$W(u,v) = \frac{H^*(u,v)}{|H(u,v)|^2 + K(u,v)}$$

Schematically ...



Restoration with a Wiener filter

$$G(u,v) = H(u,v) F(u,v) + N(u,v)$$

$$\hat{F}(u,v) = W(u,v) G(u,v)$$



# Example 1: Focus deblurring with a Wiener filter

blur  $\sigma = 1.5$  pixels

noise  $\sigma = 0.3$  grey levels

$$\hat{F}(u,v) = W(u,v) G(u,v)$$

$$W(u,v) = \frac{H^*(u,v)}{|H(u,v)|^2 + K(u,v)}$$

$g(x,y)$



$\hat{f}(x,y)$



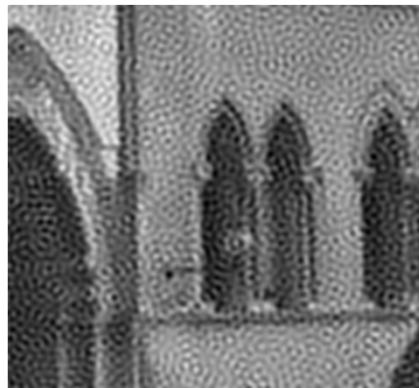
$K = 1.0 \text{ e } -5$



$K = 1.0 \text{ e } -3$



$K = 1.0 \text{ e } -1$



blur  $\sigma = 3.0$  pixels

noise  $\sigma = 0.3$  grey levels

$f(x,y)$



$g(x,y)$



$\hat{f}(x,y)$



$K = 5.0 \text{ e } -4$



# Wiener filter – sketch derivation

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Aim is to find filter which minimizes

$$\mathcal{E} = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left( f(x, y) - \hat{f}(x, y) \right)^2 dx dy$$

$$\begin{aligned} \mathcal{E} &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left| f(x, y) - \hat{f}(x, y) \right|^2 dx dy \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left| F(u, v) - \hat{F}(u, v) \right|^2 du dv \end{aligned} \quad \text{Parseval's Theorem}$$

$$\hat{F} = WG = WHF + WN$$

$$F - \hat{F} = (1 - WH)F - WN$$

$$\begin{aligned} \mathcal{E} &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} |(1 - WH)F - WN|^2 du dv \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left\{ |(1 - WH)F|^2 + |WN|^2 \right\} du dv \end{aligned} \quad \text{since } f(x, y) \text{ and } \eta(x, y) \text{ uncorrelated}$$

- Note, integrand is sum of two squares

Minimize integral if integrand minimum for all  $(u,v)$

NB  $\frac{\partial}{\partial z}(zz^*) = 2z^*$

$$\frac{\partial}{\partial z} \rightarrow 2 \left( -(1 - W^* H^*) H |F|^2 + W^* |N|^2 \right) = 0$$

$$W^* = \frac{H |F|^2}{|H|^2 |F|^2 + |N|^2}$$

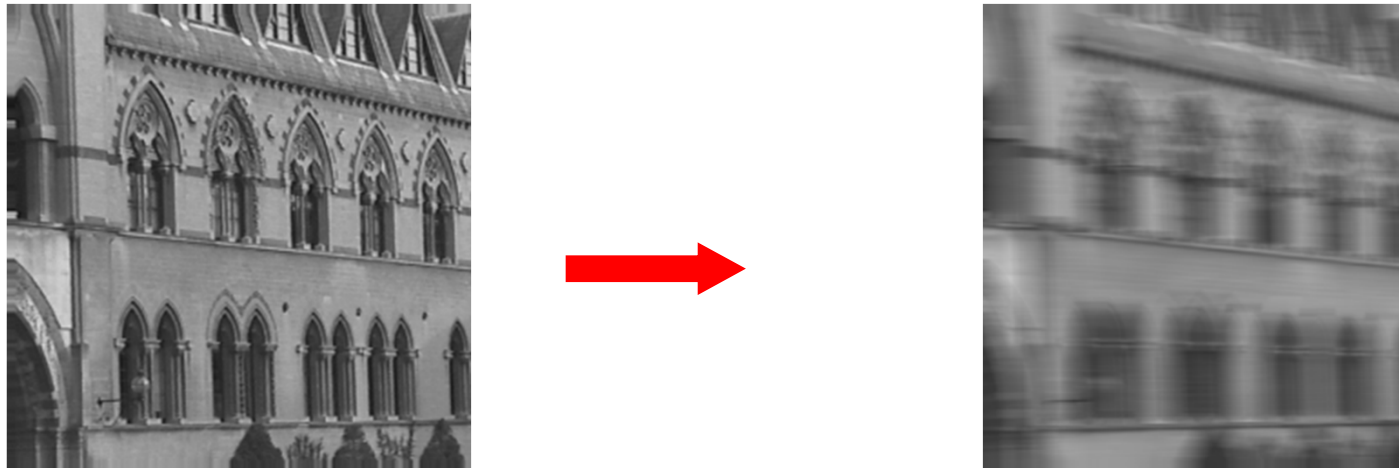
$$W = \frac{H^*}{|H|^2 + |N|^2 / |F|^2}$$

Note: filter is defined in the Fourier domain

## Example 2: Motion deblurring

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Suppose there is blur only in the horizontal direction  
e.g. camera pans as image is acquired



Degradation model

$$g(x, y) = \frac{1}{T} \int_{-T/2}^{T/2} f(x - x_0(t), y) dt$$

Require  $H(u, v)$  for Wiener filter

$$\begin{aligned}
 G(u, v) &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(x, y) e^{-j2\pi(ux+vy)} dx dy, \\
 &= \frac{1}{T} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left\{ \int_{-T/2}^{T/2} f(x - x_0(t), y) dt \right\} e^{-j2\pi(ux+vy)} dx dy
 \end{aligned}$$

interchange order of spatial and temporal integration

$$G(u, v) = \frac{1}{T} \int_{-T/2}^{T/2} \underbrace{\left\{ \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(x - x_0(t), y) e^{-j2\pi(ux+vy)} dx dy \right\}}_{\text{Fourier transform of } f(x-x_0(t), y)} dt$$

$$\begin{aligned}
 G(u, v) &= \frac{1}{T} \int_{-T/2}^{T/2} F(u, v) e^{-j2\pi u x_0(t)} dt \\
 &= F(u, v) \frac{1}{T} \int_{-T/2}^{T/2} e^{-j2\pi u x_0(t)} dt \\
 &= F(u, v) H(u, v)
 \end{aligned}$$

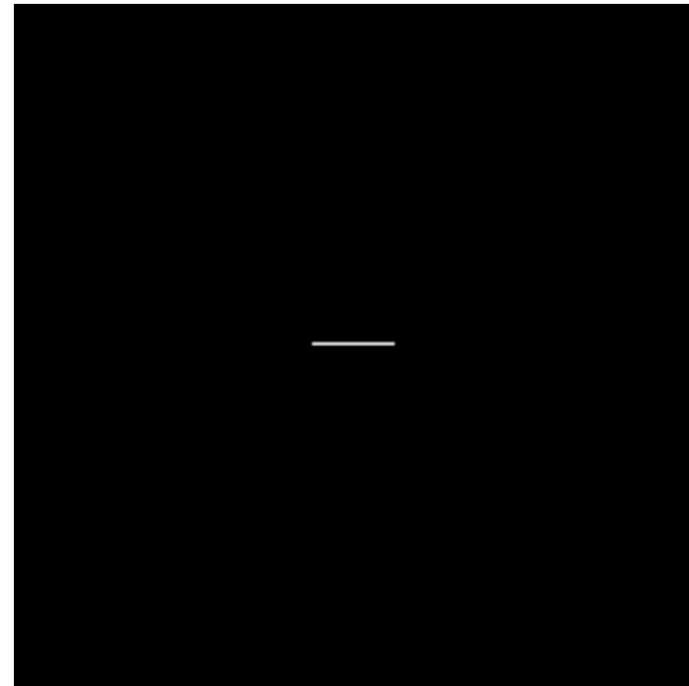
where

$$H(u, v) = \frac{1}{T} \int_{-T/2}^{T/2} e^{-j2\pi u x_0(t)} dt$$

suppose  $x_0(t) = st$ , and  $sT = d$  pixels

FT of ...  $h(x, y) = \text{hat}_d(x)\delta(y)$

$$\begin{aligned} H(u, v) &= \frac{1}{T} \int_{-T/2}^{T/2} e^{-j2\pi ust} dt \\ &= \frac{1}{T} \left[ \frac{e^{-j2\pi ust}}{-j2\pi us} \right]_{-T/2}^{T/2} \\ &= \frac{1}{j2\pi ud} (e^{j\pi ud} - e^{-j\pi ud}) \\ &= \text{sinc}\pi ud \end{aligned}$$



Note,  $H(u, v)$  has zeros – a problem for an inverse filter

# Motion deblurring with a Wiener filter

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blur = 20 pixels

$$W(u, v) = \frac{H^*(u, v)}{|H(u, v)|^2 + K(u, v)}$$



1. Compute the FT of the blurred image
2. Multiply the FT by the Wiener filter
3. Compute the inverse FT

$$\hat{F}(u, v) = W(u, v) G(u, v)$$

# Application: Reading number plates

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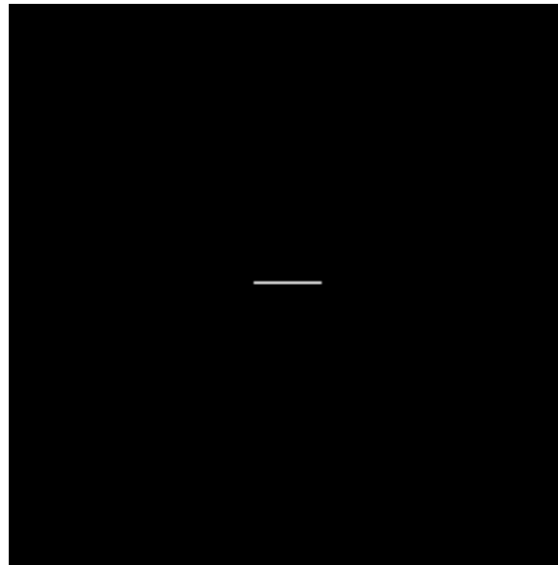
## Algorithm

1. Rotate image so that blur is horizontal
2. Estimate length of blur
3. Construct a bar modelling the convolution
4. Compute and apply a Wiener filter
5. Optimize over values of  $K$

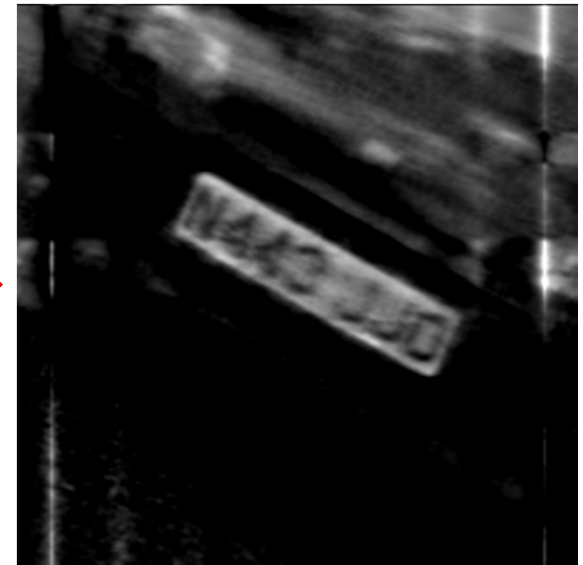
$f(x,y)$



$h(x,y)$



$\hat{f}(x,y)$



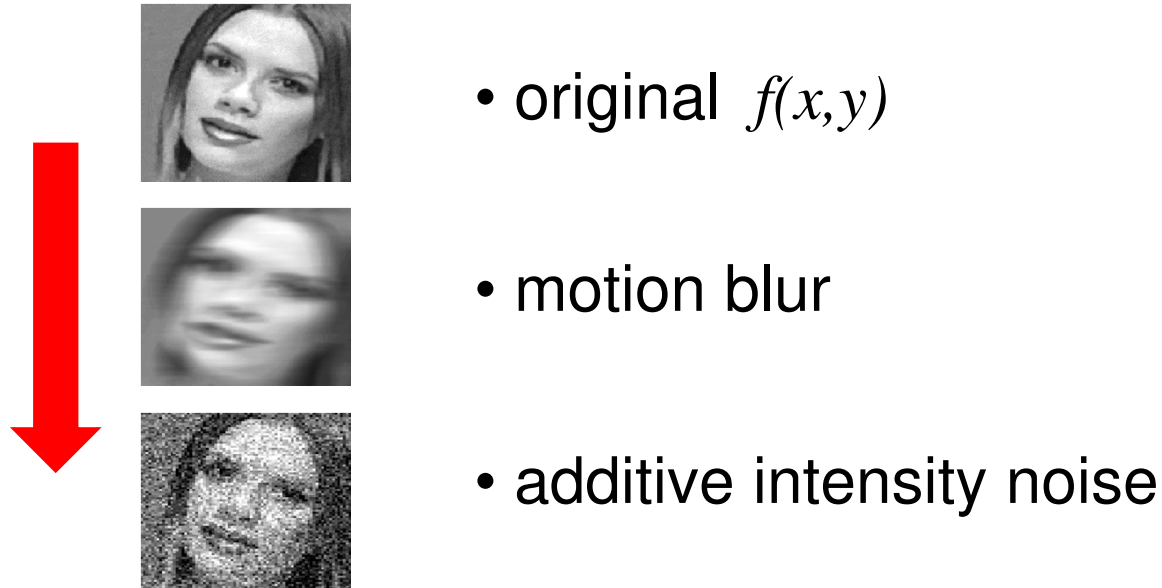
blur = 30 pixels



# Maximum a posteriori (MAP) Estimation

# Generative model (forward process)

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For an image with  $n$  pixels, write this process as

$$\hat{\mathbf{g}} = \mathbf{A}\mathbf{f} + \mathbf{n}$$

where  $\hat{\mathbf{g}}$  and  $\mathbf{f}$  are  $n$ -vectors, and  $\mathbf{A}$  is an  $n \times n$  matrix.

# Inverse problem

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- Estimate  $f(x,y)$  by optimizing a cost function:

$$\hat{\mathbf{f}} = \arg \min_{\mathbf{f}} \underbrace{(\underbrace{\mathbf{g}}_{\substack{\text{observed} \\ \text{image}}} - \underbrace{\mathbf{A}\mathbf{f}}_{\substack{\text{generated} \\ \text{image}}})^2}_{\substack{\text{Likelihood/} \\ \text{loss function}}} + \underbrace{\lambda p(\mathbf{f})}_{\substack{\text{prior/} \\ \text{regularization}}}$$

Example

$$p(f) = (\nabla \mathbf{f})^2$$

to suppress high frequency noise

## Example 3: Super resolution

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Suppose there are multiple images of the same scene each displaced spatially ...

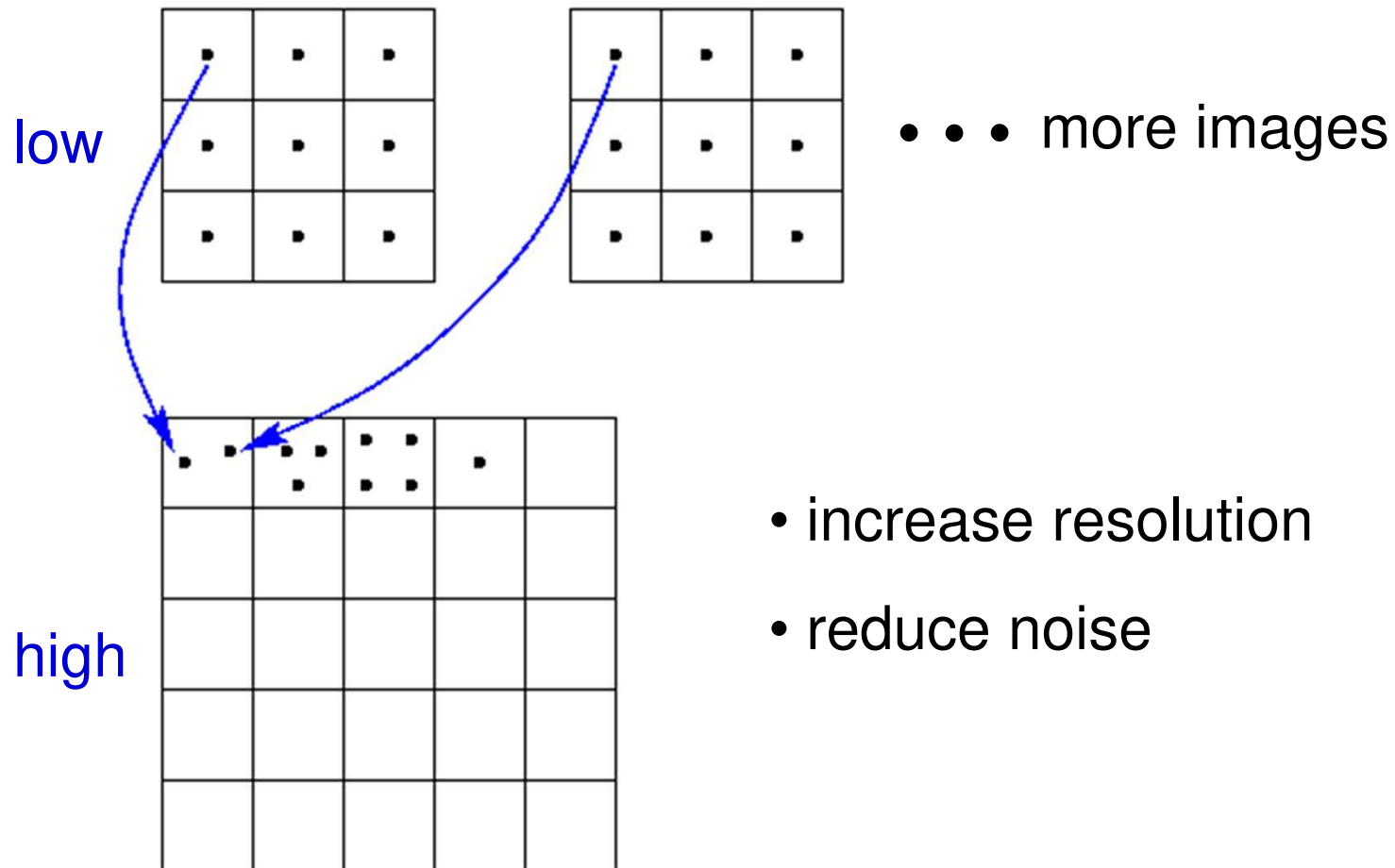


After registration the samples are not coincident and this may be used to defeat the Nyquist limit.

# Intuitive model

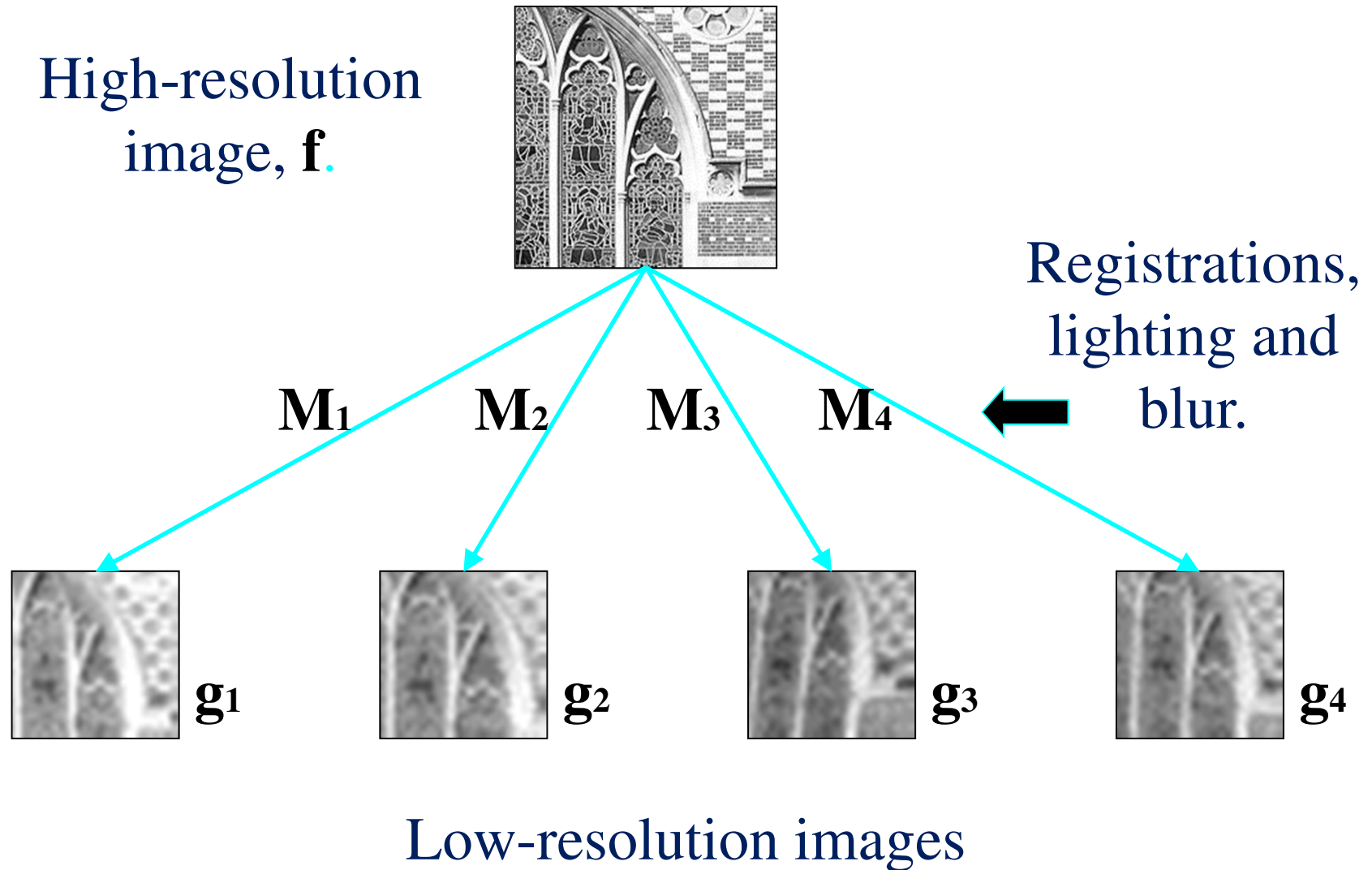
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Treat images as point samples



# Generative Model

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## Sketch solution

Non-examinable

- Estimate the super resolution image which minimizes the error between predicted and observed images.

Write the generative model for one image  $i$  as

$$\mathbf{g}_i = \mathbf{M}_i \mathbf{f} + \boldsymbol{\eta}_i$$

where  $\mathbf{M}_i$  combines registration, lighting and down-sampling. Finally the generative models of all  $N$  images are stacked vertically to form an over-determined linear system

$$\begin{bmatrix} \mathbf{g}_0 \\ \mathbf{g}_1 \\ \vdots \\ \mathbf{g}_{N-1} \end{bmatrix} = \begin{bmatrix} \mathbf{M}_0 \\ \mathbf{M}_1 \\ \vdots \\ \mathbf{M}_{N-1} \end{bmatrix} \mathbf{f} + \begin{bmatrix} \boldsymbol{\eta}_0 \\ \boldsymbol{\eta}_1 \\ \vdots \\ \boldsymbol{\eta}_{N-1} \end{bmatrix}$$

$$\mathbf{g} = \mathbf{M} \mathbf{f} + \boldsymbol{\eta}$$

# Maximum a posterior estimation

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The MAP estimator has the form:

$$\mathbf{f}_{\text{MAP}} = \underset{\mathbf{f}}{\operatorname{argmin}} \left( \underbrace{(\mathbf{g} - \mathbf{M}\mathbf{f})^2}_{\text{likelihood}} + \underbrace{\lambda^2 p(\mathbf{f})}_{\text{prior}} \right)$$

Example

$$\mathbf{f}_{\text{MAP}} = \underset{\mathbf{f}}{\operatorname{argmin}} \left( (\mathbf{g} - \mathbf{M}\mathbf{f})^2 + \lambda^2 \sum_{\forall x,y} p(\nabla f(x,y)) \right)$$

where the prior  $p(x)$  is defined by the [Huber function](#),

$$\begin{aligned} p(x) &= x^2, \text{ if } x \leq \alpha \\ &= 2\alpha |x| - \alpha^2, \text{ otherwise} \end{aligned}$$

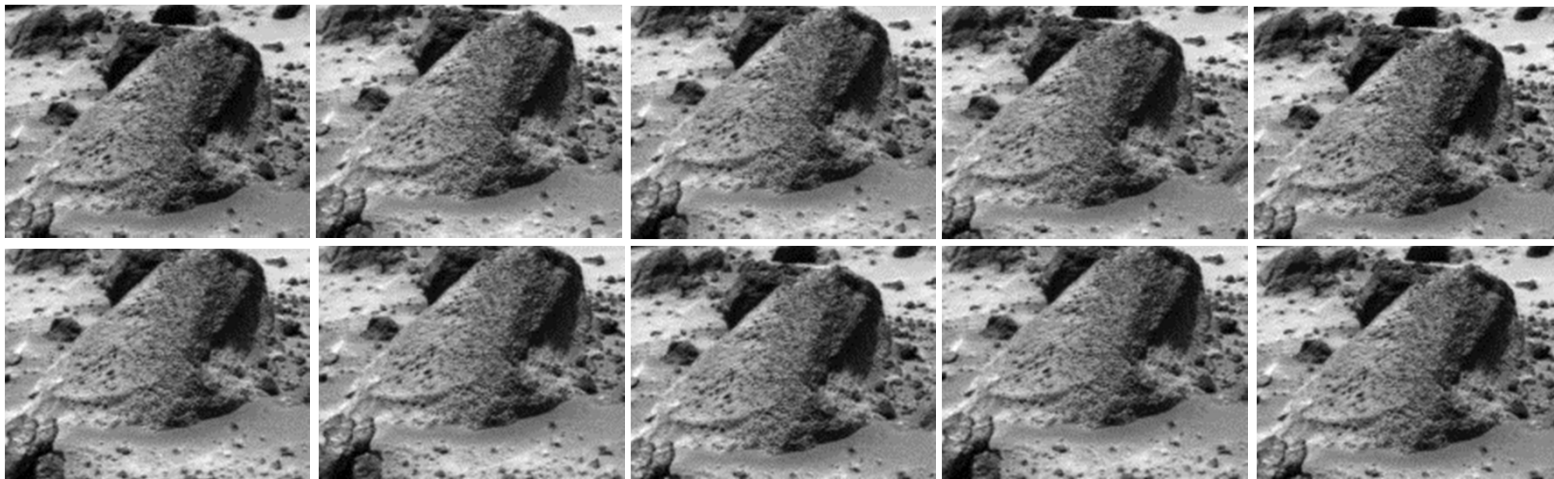


# Super resolution example I: Mars

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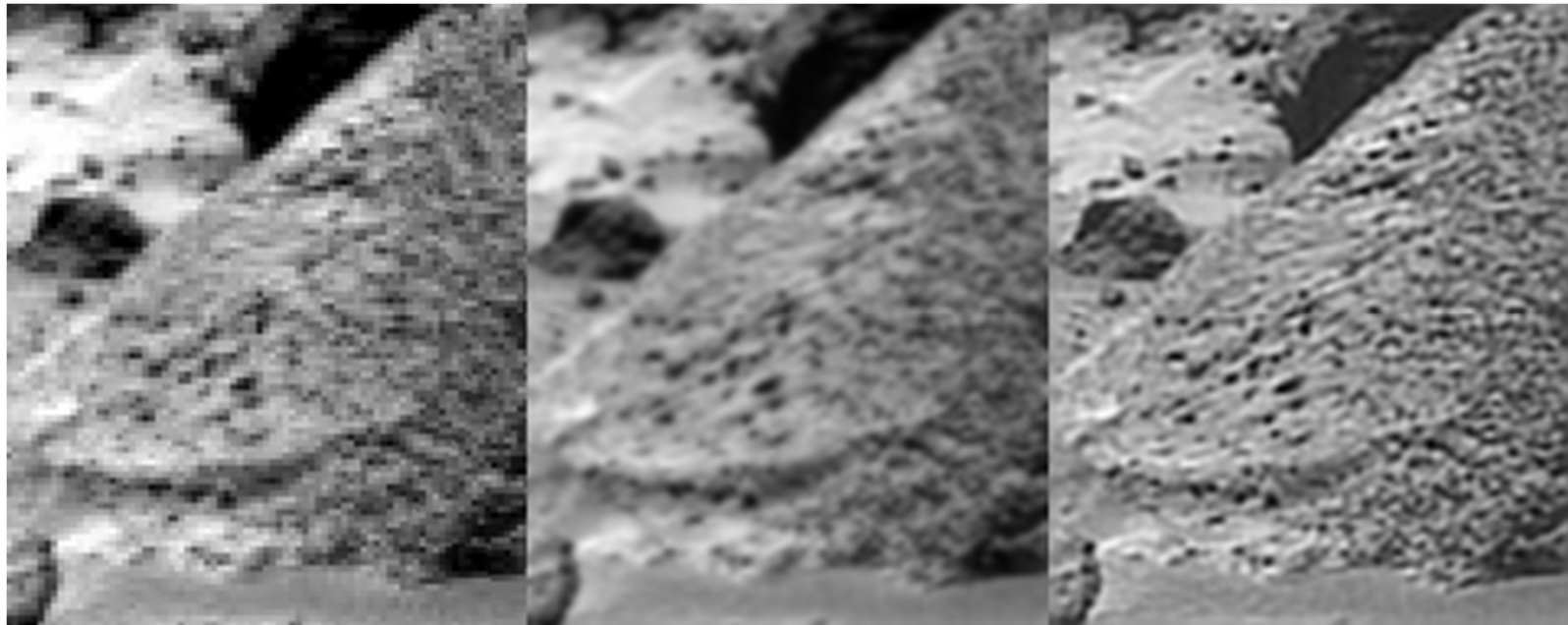
25 JPEG images courtesy of the Mars lander

images are from different sweeps of a rotating camera



# Super resolution result

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Original frame

Average image

Super-resolution

2x zoom from 25 JPEG images.

# Super resolution example II: car sequence

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rotating DV camera



# Mosaic

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# Super-resolution result for ROI

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85 JPEG images



original ROI

35 x 20 pixels



four times resolution

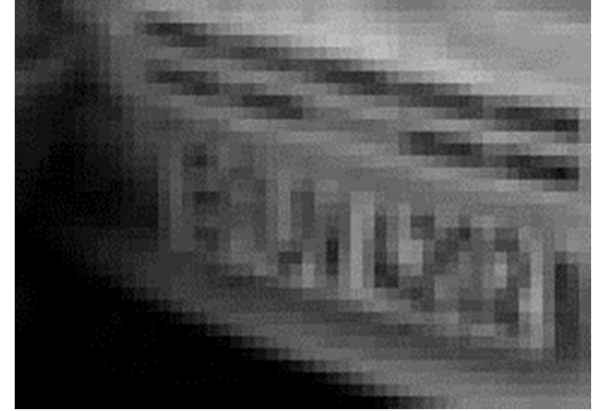
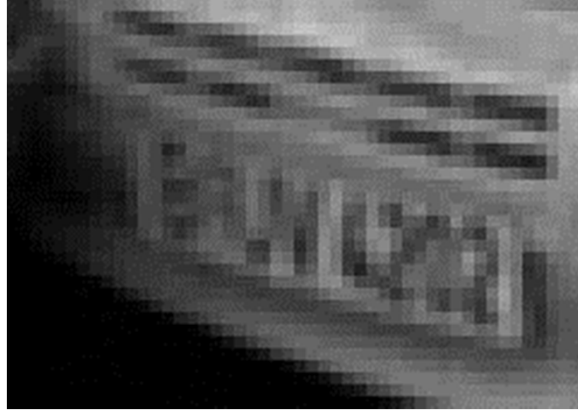
## Super resolution example III: Run Lola Run

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# Input – low resolution

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# Super-resolution output

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# Blind deblurring

Non-examinable

So far we have assumed that we know the generative model, e.g.

$$g = A(h) f$$

$$G = H F$$



i.e. that  $h(x,y)$  is known, so that given the observed image  $g(x,y)$ , then the original image  $f(x,y)$  can be estimated (restored)

Consider if only the observed image  $g(x,y)$  is known. This is the problem of **blind estimation**.

# Blind deblurring continued

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- Estimate  $f(x,y)$  and  $h(x,y)$  by optimizing a cost function:

$$\min_{\mathbf{f}, \mathbf{h}} \underbrace{(\underbrace{\mathbf{g}}_{\text{observed image}} - \underbrace{\mathbf{A}(\mathbf{h}) \mathbf{f}}_{\text{generated image}})^2}_{\text{Likelihood/loss function}} + \underbrace{\lambda p_f(\mathbf{f})}_{\text{image prior}} + \underbrace{\mu p_h(\mathbf{h})}_{\text{blur prior}}$$

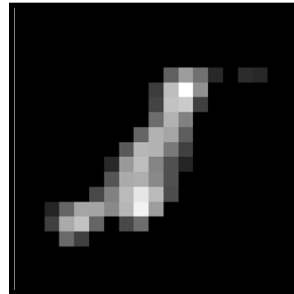
# Example I: Blind deblurring

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blurred image



estimated  
blur filter



restored image



## More examples of blind deblurring

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